



# Electronics – 96032

 POLITECNICO DI MILANO



## Amplitude Modulation and Synchronous Detection

Alessandro Spinelli

Phone: (02 2399) 4001

alessandro.spinelli@polimi.it

spinelli.faculty.polimi.it

# Disclaimer

Slides are supplementary  
material and are NOT a  
replacement for textbooks  
and/or lecture notes

# Purpose of the lesson

- It's time to begin a discussion on the techniques for improving S/N
- Noise-reduction techniques obviously depend on the type of signal and of noise:

Signal	Noise	
	HF (White)	LF (flicker)
LF (constant)	done	this lesson
HF (pulse)	done	done

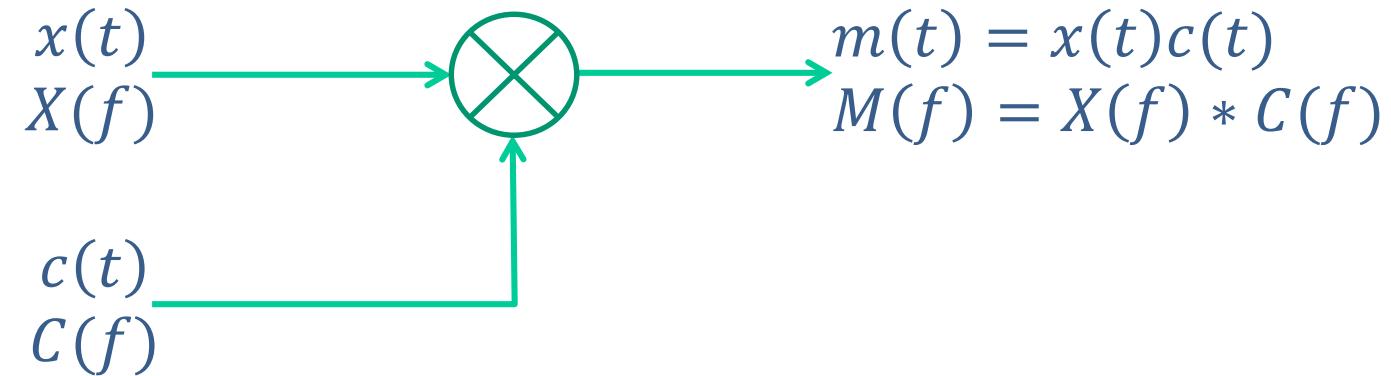
# Outline

- Amplitude modulation and demodulation
- The weighting function

# The problem

- If the signal is at DC level and buried in LF noise, HP or BP filters become useless
- If we could move the signal spectrum to higher frequencies, we would improve  $S/N$
- A high-Q BP filter could then be used to recover the signal

# Amplitude modulation (AM)



- The amplitude of a carrier wave  $c(t)$  is modified by a modulating signal  $x(t)$
- Originally developed for telephone and radio communication

# Time and frequency domains

- We consider a sinusoidal carrier

$$c(t) = A \cos(\omega_c t + \phi_c) \Leftrightarrow$$

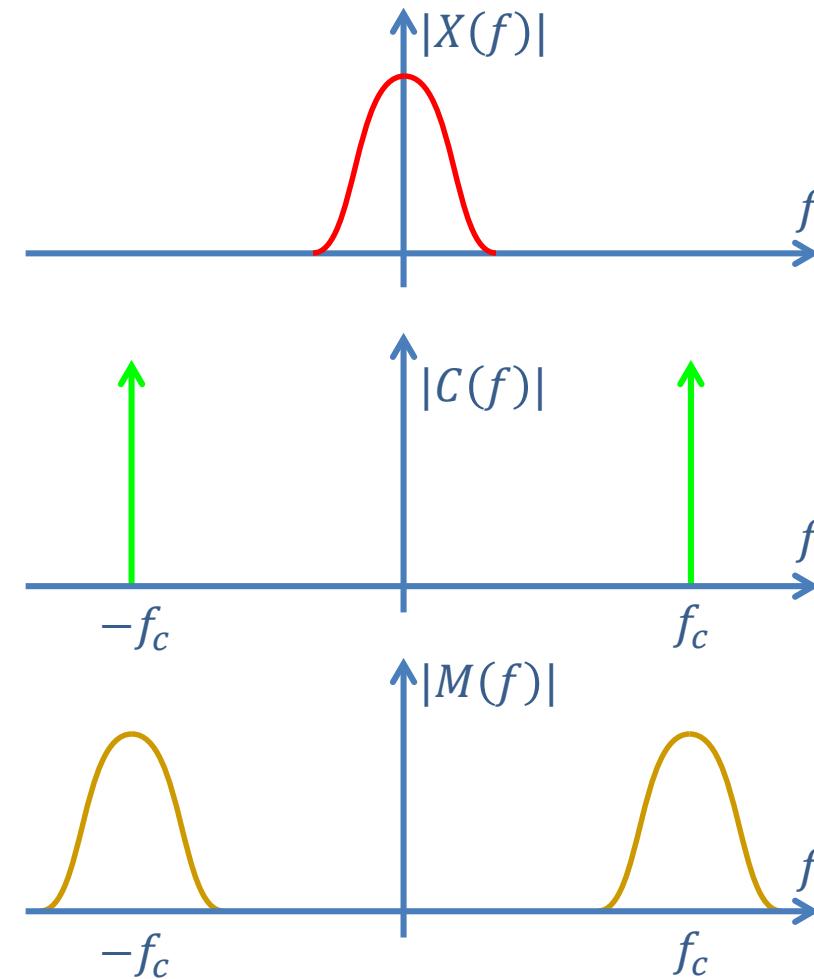
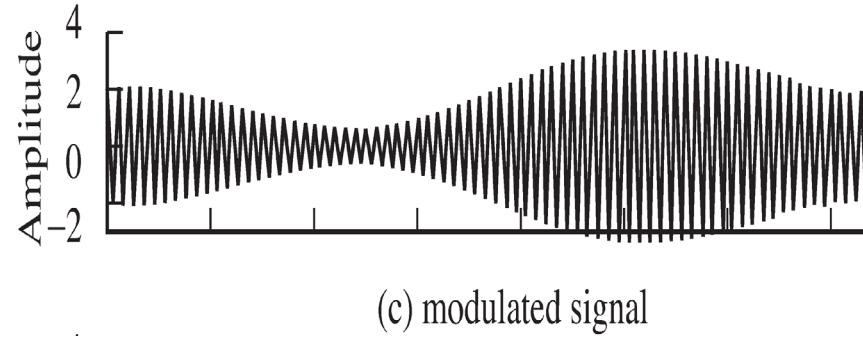
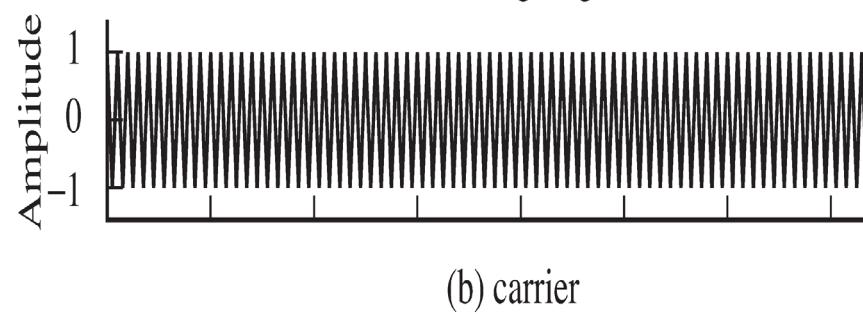
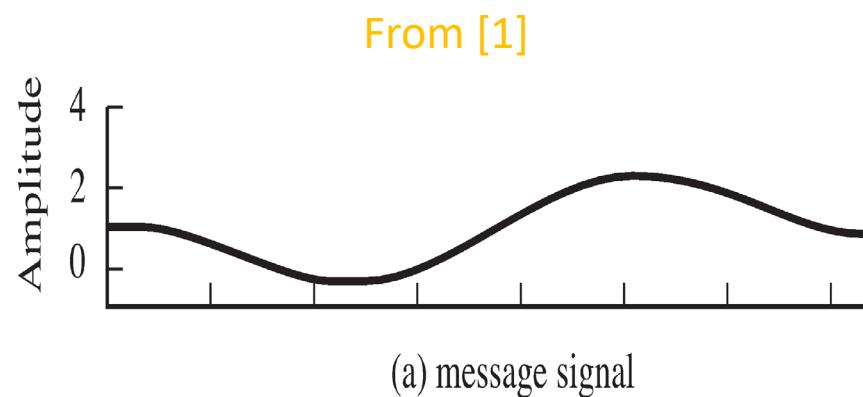
$$C(f) = \frac{A}{2} \left( e^{j\phi_c} \delta(f - f_c) + e^{-j\phi_c} \delta(f + f_c) \right)$$

- The modulated signal is then

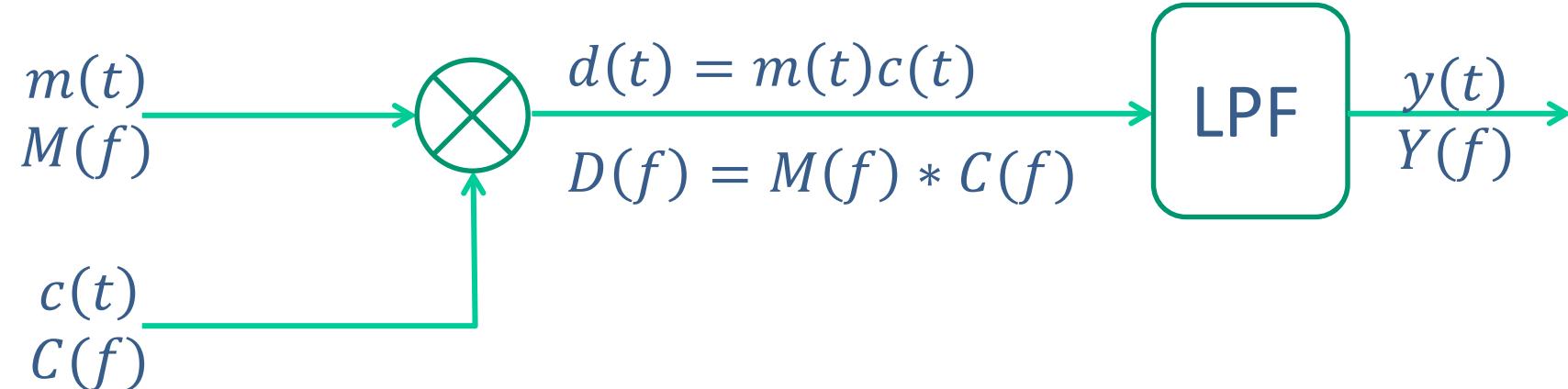
$$m(t) = x(t)c(t) \Leftrightarrow$$

$$M(f) = X(f) * C(f) = \frac{A}{2} \left( e^{j\phi_c} X(f - f_c) + e^{-j\phi_c} X(f + f_c) \right)$$

# Waveforms and spectra



# Demodulation

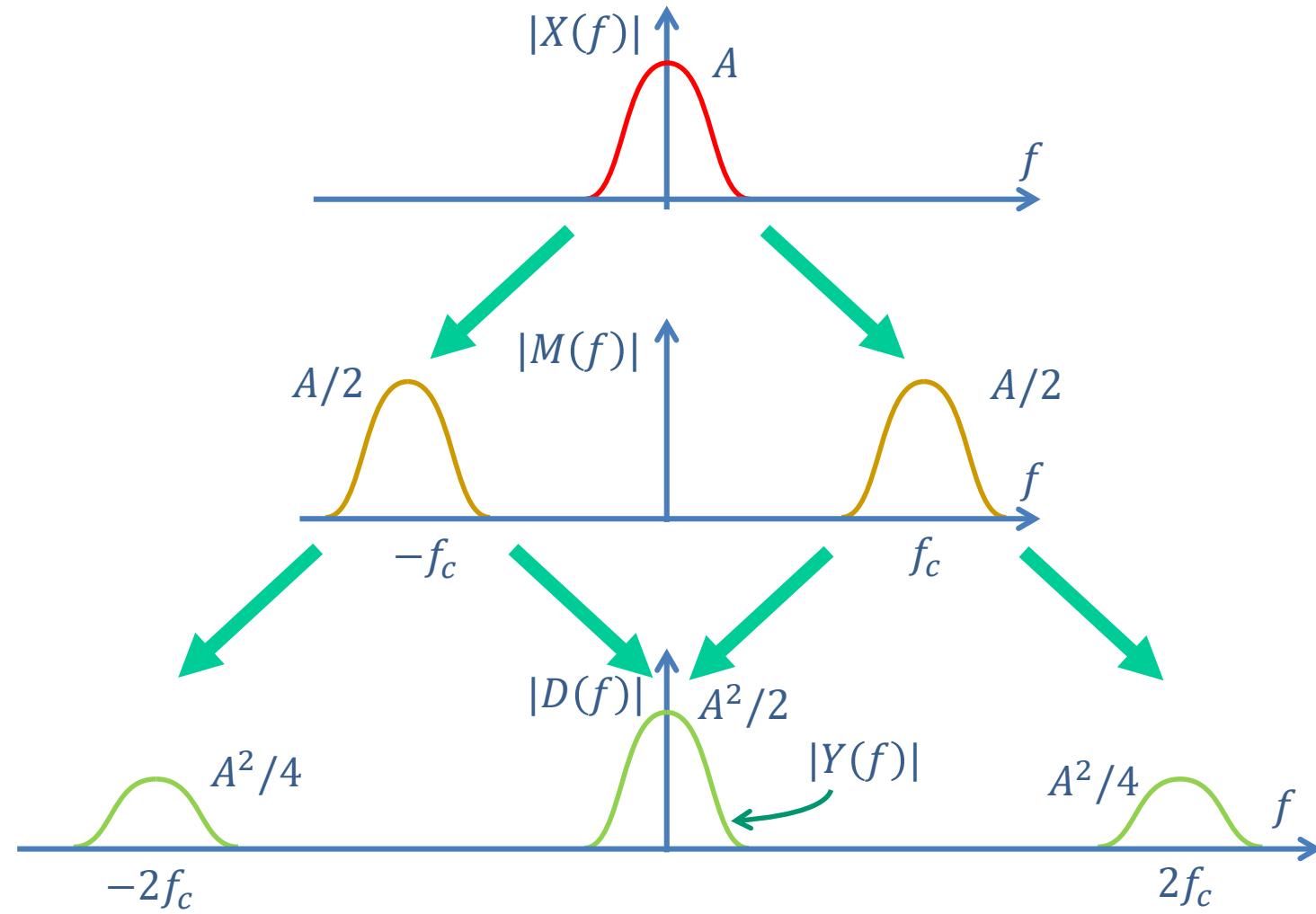


$$d(t) = m(t)c(t) = x(t)c^2(t) = x(t)A^2\cos^2(\omega_c t + \phi_c)$$

$$Y(f) = \frac{A^2}{2}x(t)(1 + \cos(2\omega_c t + 2\phi_c))$$

$$D(f) = \frac{A^2}{2}X(f) + \frac{A^2}{4}\left(e^{j2\phi_c}X(f - 2f_c) + e^{-j2\phi_c}X(f + 2f_c)\right)$$

# Frequency-domain view



# Frequency and phase errors

- We consider demodulating with

$$w_R(t) = B \cos((\omega_c + \Delta\omega)t + \phi_c + \Delta\phi)$$

- The low-frequency term (small  $\Delta\omega$ ) is

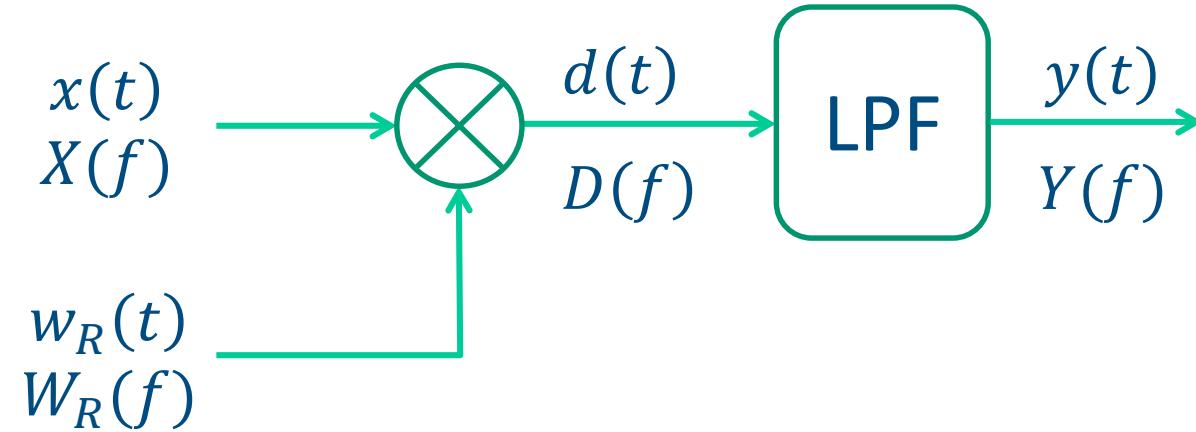
$$y(t) = ABx(t) \cos(\Delta\omega t + \Delta\phi)/2$$

- Phase error  $\Rightarrow$  signal reduction
- Frequency error  $\Rightarrow$  oscillating behavior
- The reference must be locked in frequency and phase to the carrier  $\Rightarrow$  **synchronous detection**
- The demodulator is also called **phase-sensitive detector (PSD)**

# Outline

- Amplitude modulation and demodulation
- The weighting function

# PSD weighting function

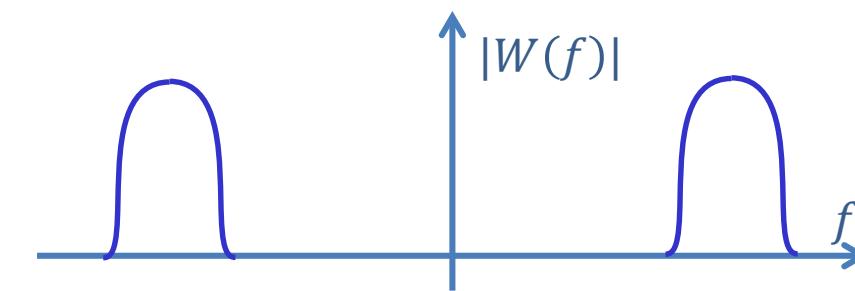
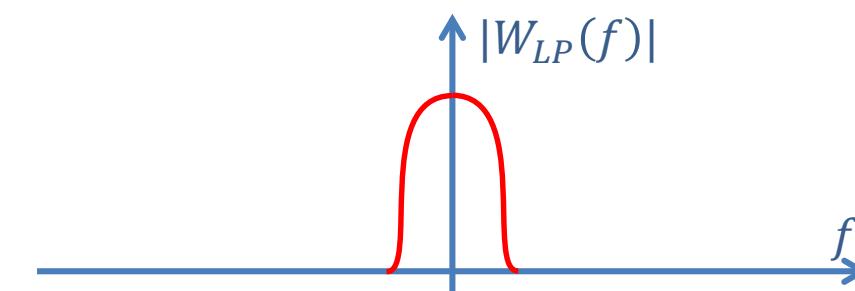
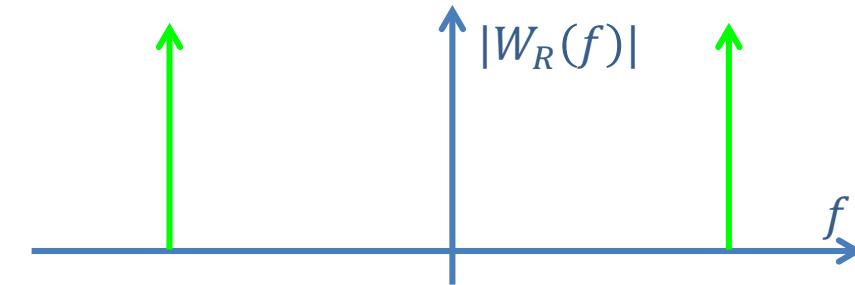
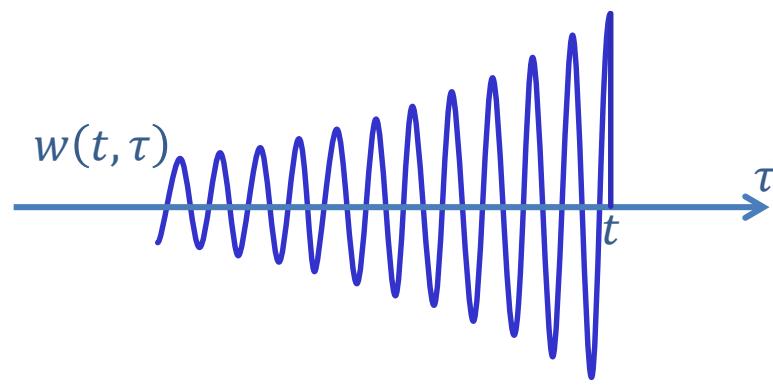
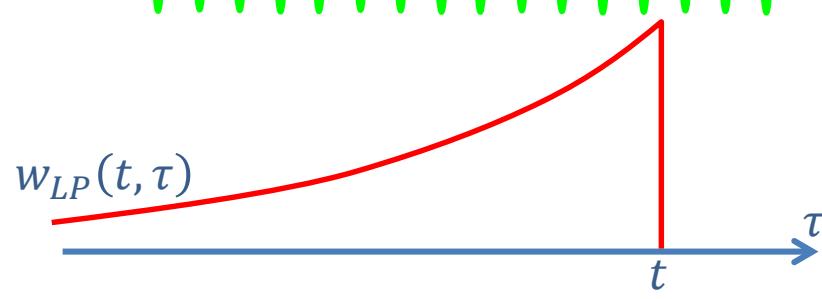
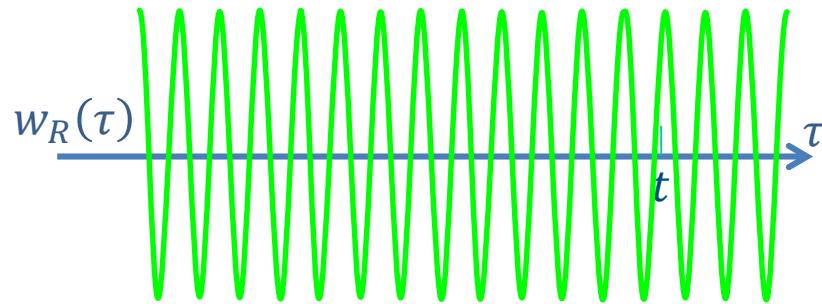


$$y(t) = \int d(\tau)w_{LP}(t, \tau)d\tau = \int x(\tau)w_R(\tau)w_{LP}(t, \tau)d\tau$$

$$w(t, \tau) = w_R(\tau)w_{LP}(t, \tau)$$

$$W(t, f) = W_R(f) * W_{LP}(t, f)$$

# Time/frequency domains



- The filter is time-variant even if LPF is LTI  $\Rightarrow$  remember that

$$y(t) = \int x(\tau)w(t, \tau)d\tau = \int X(f)W^*(t, f)df$$

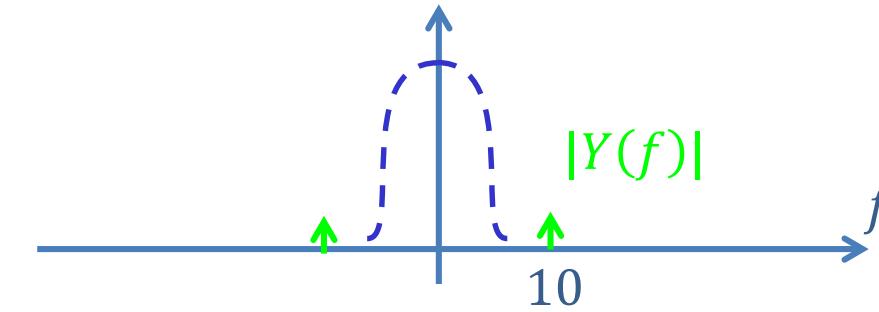
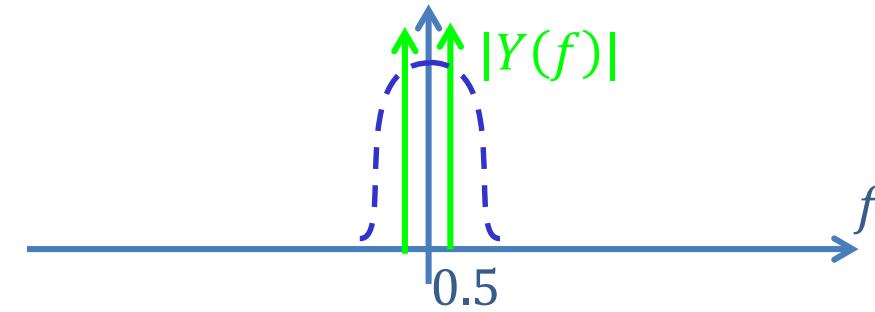
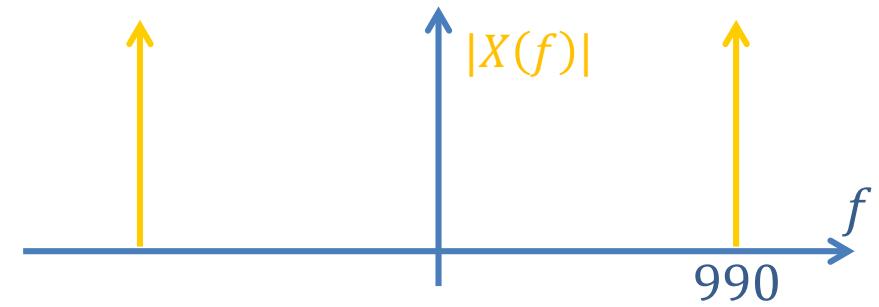
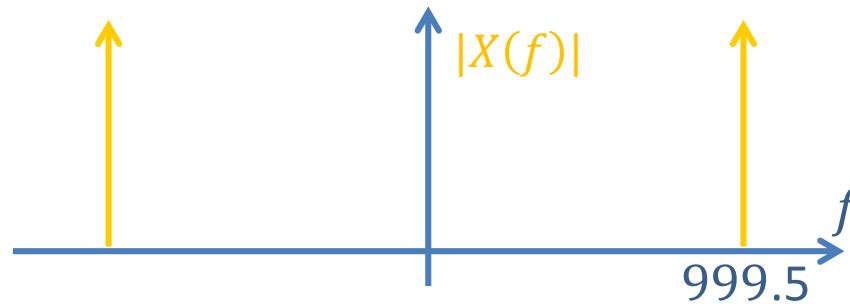
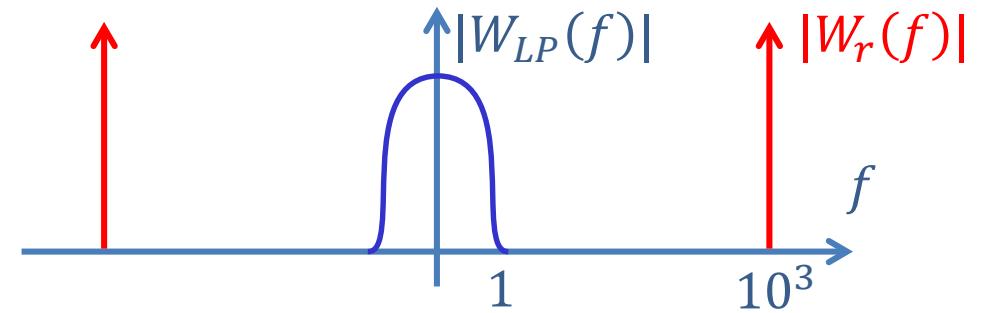
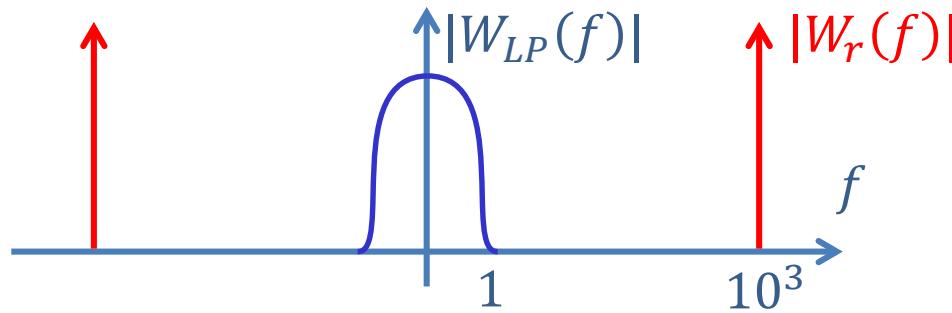
$\Rightarrow Y(f)$  is NOT equal to  $X(f)W(t, f)$

- For the case of a PSD with LTI LPF, we have instead

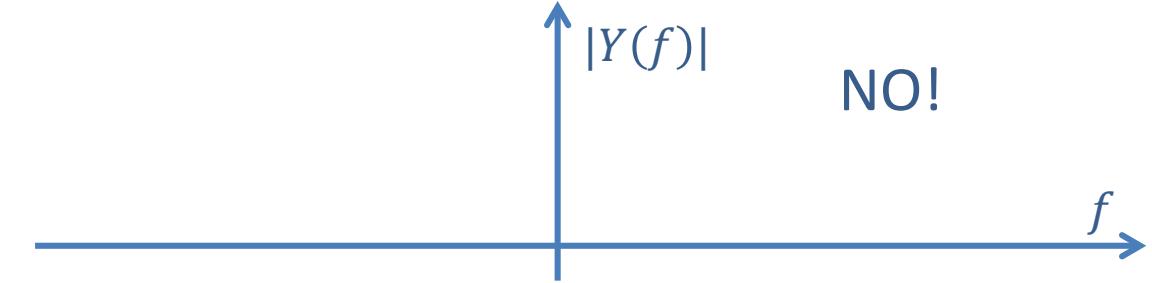
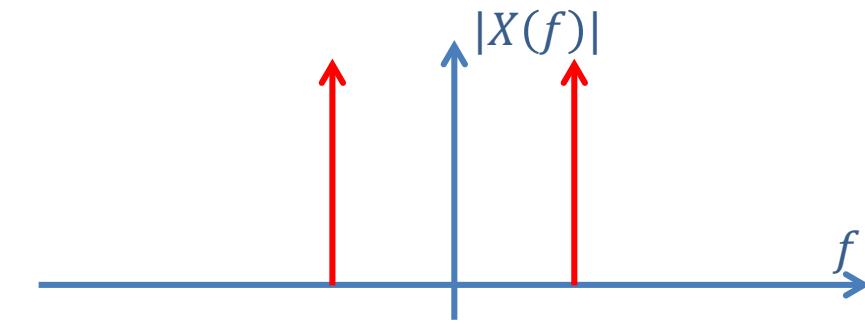
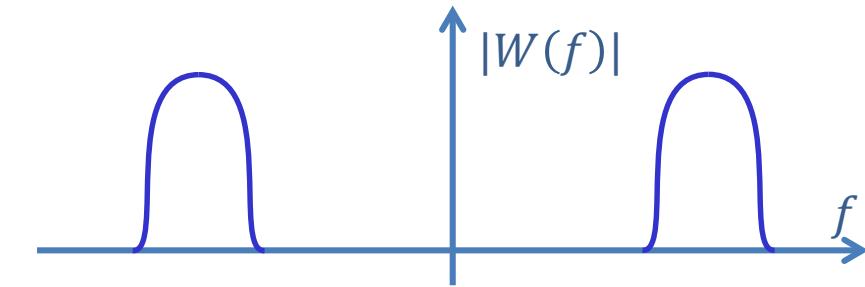
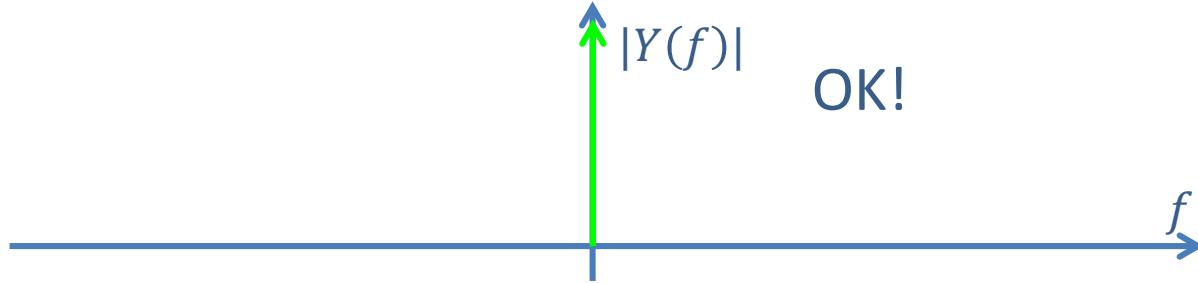
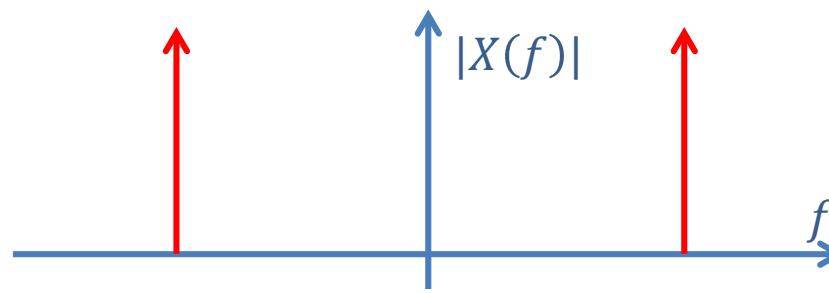
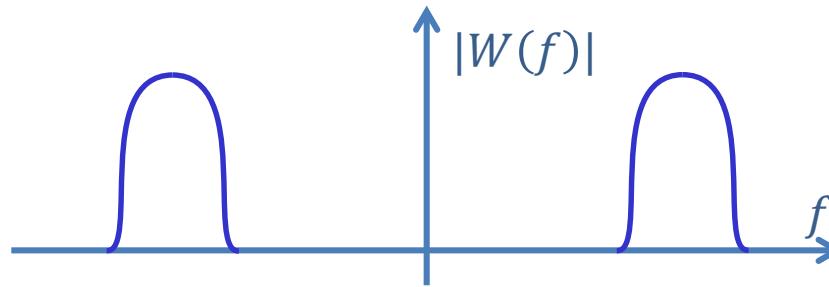
$$Y(f) = H_{LP}(f)(X(f) * W_R(f))$$

- If  $w_R(t)$  is a periodic signal,  $|W(f)|$  represents the frequency components that give contributions in the baseband

# Examples



# Weighting function interpretation



# PSD as optimum filter

- Let's take a simple case of constant signal  $\Rightarrow$  the input of the PSD is then  $x(t) = A\cos(\omega_r t)$
- If LPF is an LTI integrator

$$w_{LP}(t, \tau) = K u(t - \tau) \Rightarrow w(t, \tau) \propto x(\tau)$$

$\Rightarrow$  PSD is the optimum filter!

- In the general case, the signal is not constant  $\Rightarrow$  the PSD is quasi-optimum, but more flexible
  - Reference frequency set externally
  - BW of LPF can accommodate different signals

# Another look at the weighting function

- For the case of the LTI integrator we have

$$y(t) = K \int_{-\infty}^t x(\tau) w_R(\tau) d\tau \approx K_{xw_R}(0)$$

- $y(t)$  is an **estimate of the cross-correlation between input and reference signals**  $\Rightarrow$  maximum output is achieved when  $x(t)$  has the same frequency and phase as  $w_R(t)$
- For, say, a simple RC filter we have

$$y(t) = \frac{1}{T_F} \int_{-\infty}^t x(\tau) w_R(\tau) e^{-\frac{t-\tau}{T_F}} d\tau$$

which is again  $K_{xw_R}(0)$  estimated over a time  $T_F$

# Frequency domain

- In the frequency domain we have

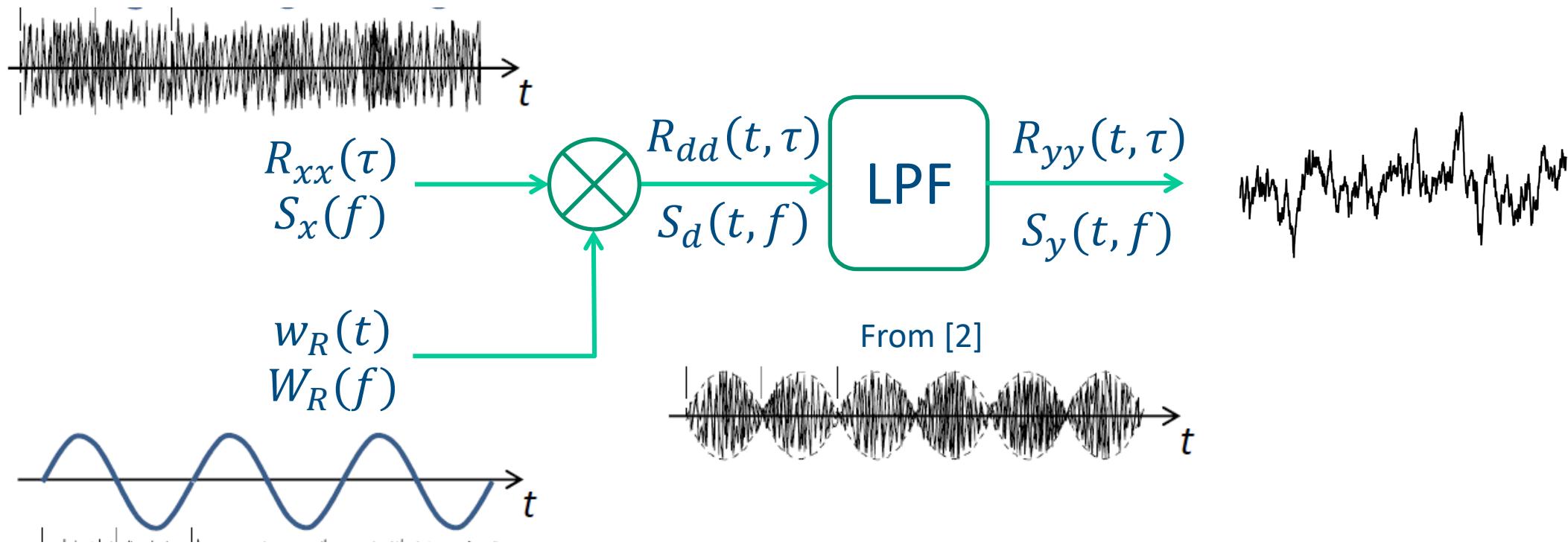
$$X(f) * W_R(f) = \int X(\nu)W_R(f - \nu)d\nu$$

- The output LPF selects the components around  $f = 0$ , i.e.

$$Y(f) \approx \int X(\nu)W_R(-\nu)d\nu = \int X(\nu)W_R^*(\nu)d\nu$$

We find again the correlation behavior!

# Output noise



The output noise of the PSD is non-stationary (actually cyclostationary) even for stationary input noise

# Output noise autocorrelation

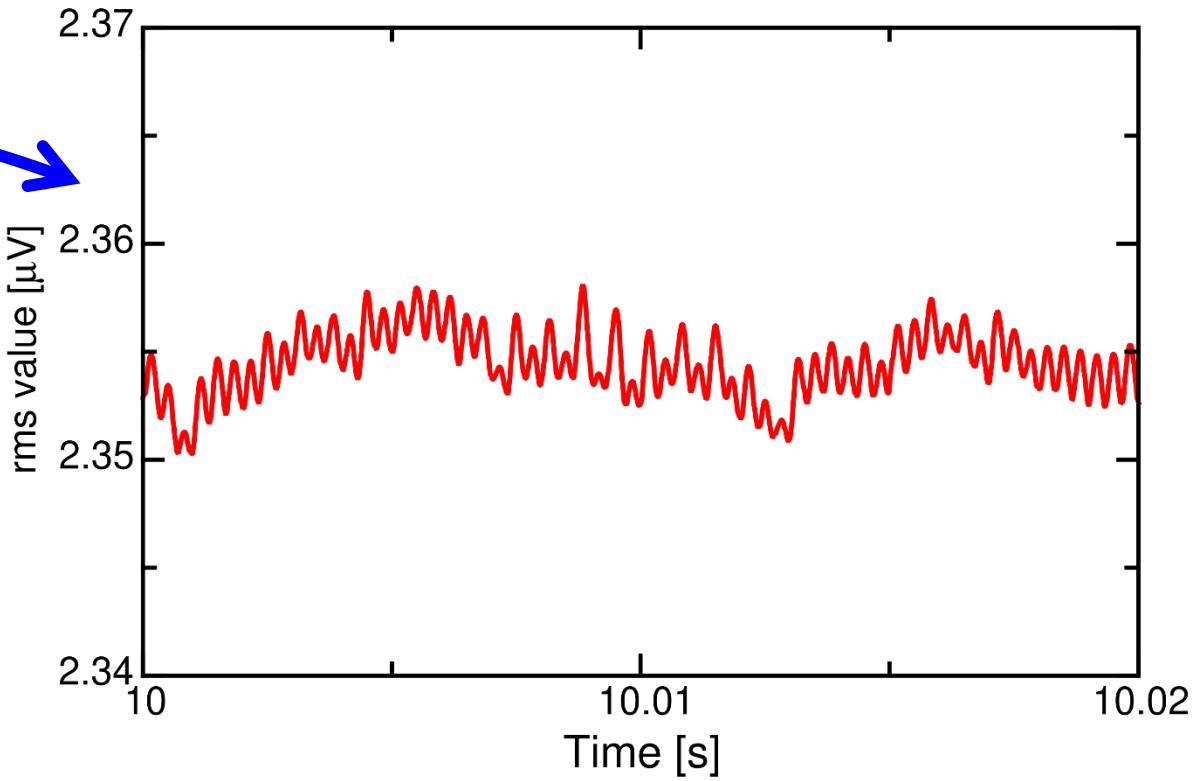
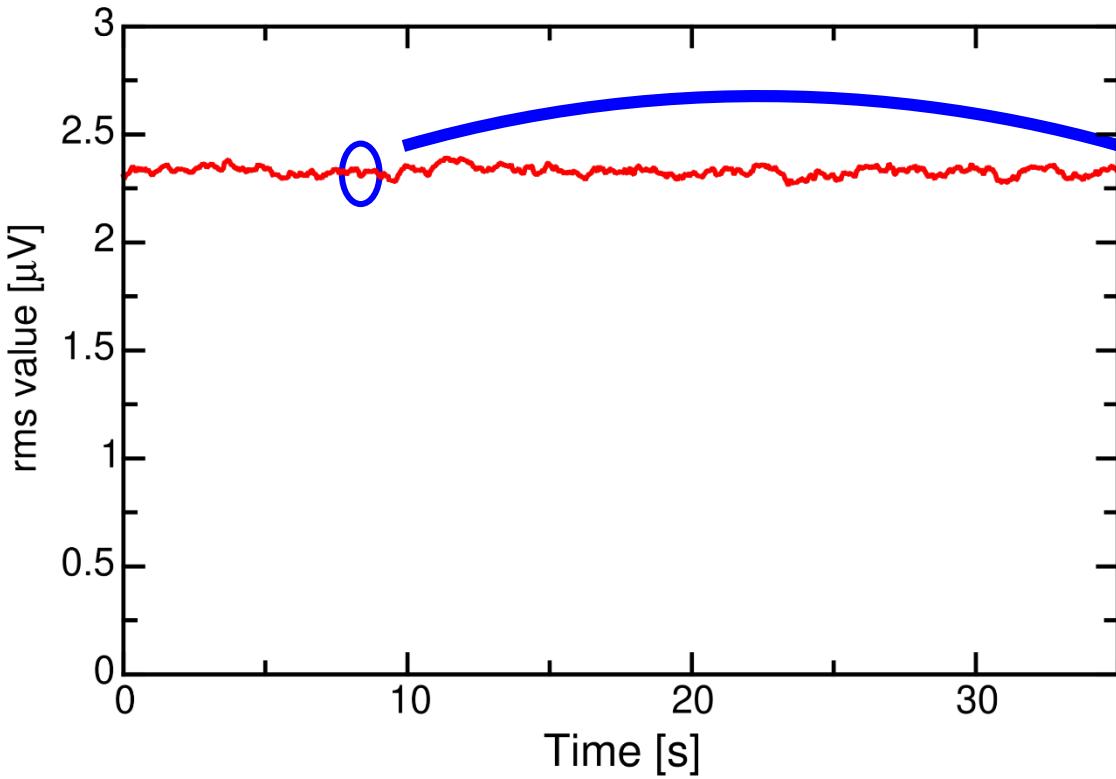
$$\begin{aligned} R_{dd}(t, t + \tau) &= \overline{n_d(t)n_d(t + \tau)} = \overline{n_x(t)n_x(t + \tau)}w_R(t)w_R(t + \tau) \\ &= R_{xx}(\tau)w_R(t)w_R(t + \tau) \end{aligned}$$

- In particular,  $\overline{n_d^2(t)} = \overline{n_x^2}w_R^2(t)$
- Since the output filter averages over many periods of  $w_R$ , we consider the time average

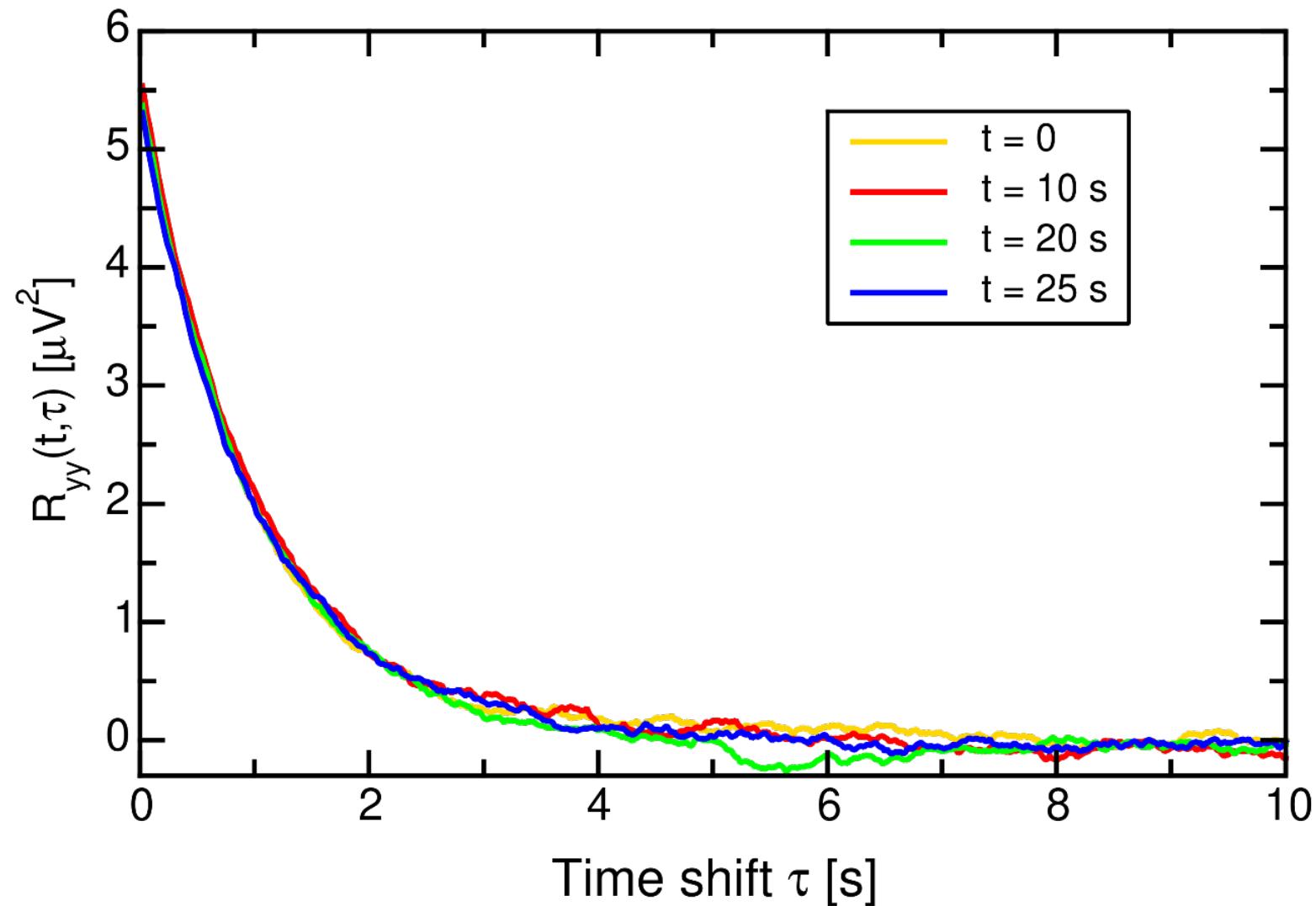
$$\begin{aligned} \langle R_{dd}(t, t + \tau) \rangle &= R_{xx}(\tau)\langle w_R(t)w_R(t + \tau) \rangle \\ &= R_{xx}(\tau) \lim_{T \rightarrow \infty} \frac{1}{2T} \int_{-T}^T w_R(t)w_R(t + \tau) dt = R_{xx}(\tau)K_{w_R w_R}(\tau) \end{aligned}$$

# Output noise rms value

- Input flicker noise,  $K = 3.3 \times 10^{-8} \text{ V}^2$ ,  $f_r = 1.5 \text{ kHz}$ ,  $T_F = 1 \text{ s}$
- Output rms noise (ensemble average over 6000 realizations):



# Output noise autocorrelation



# Wrap up

- The reference signal is

$$w_R(t) = B \cos(\omega_r t) \Leftrightarrow W_R(f) = \frac{B}{2} (\delta(f - f_r) + \delta(f + f_r))$$

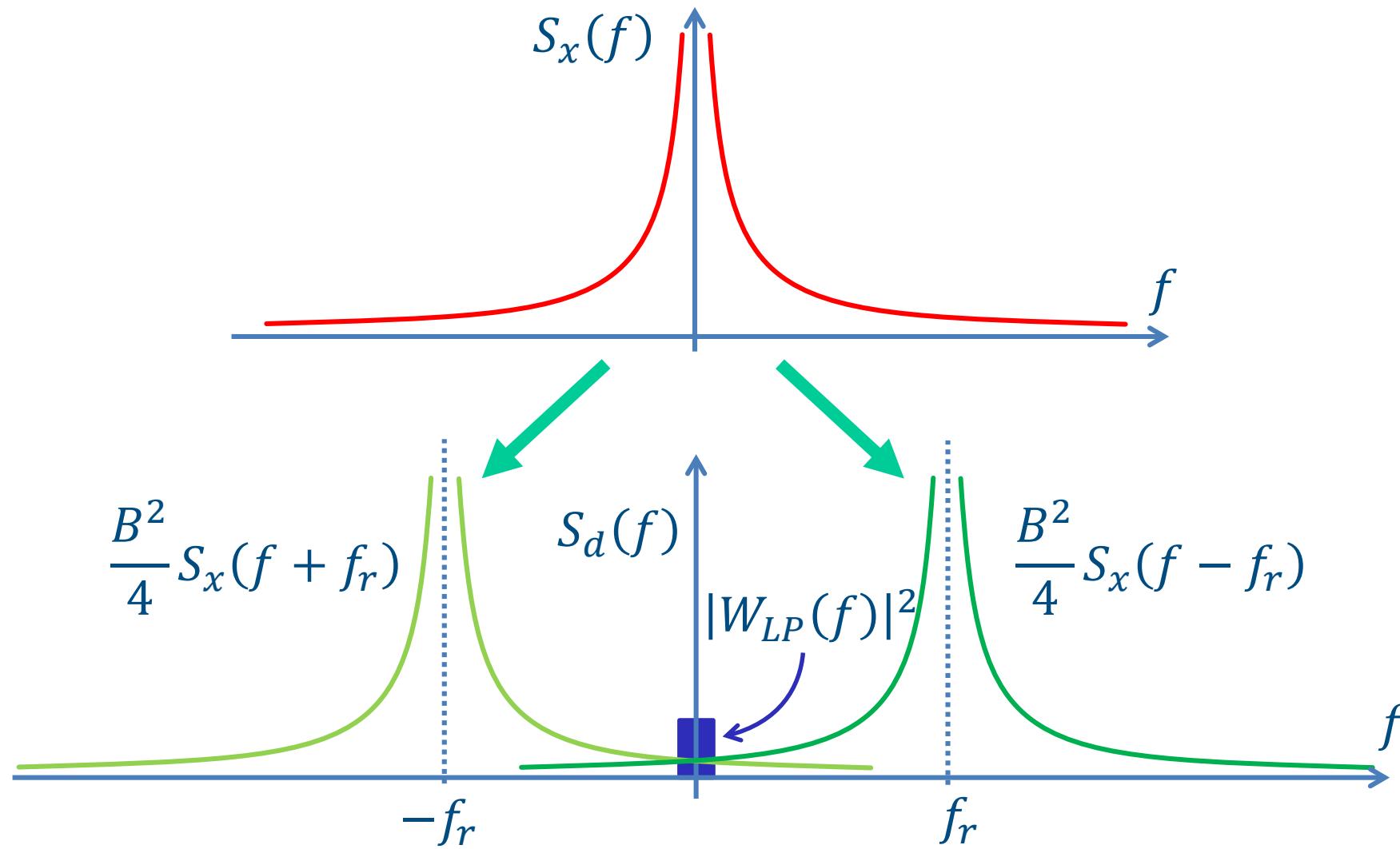
- Its time correlation is

$$K_{w_R w_R}(\tau) = \frac{B^2}{2} \cos(\omega_r \tau) \Leftrightarrow S_{w_R}(f) = \frac{B^2}{4} (\delta(f - f_r) + \delta(f + f_r))$$

- The demodulated noise autocorrelation becomes

$$R_{dd}(\tau) = R_{xx}(\tau) \frac{B^2}{2} \cos(\omega_r \tau) \Leftrightarrow S_d(f) = \frac{B^2}{4} (S_x(f - f_r) + S_x(f + f_r))$$

# Ex: flicker noise spectra (LTI LPF)



# Output noise

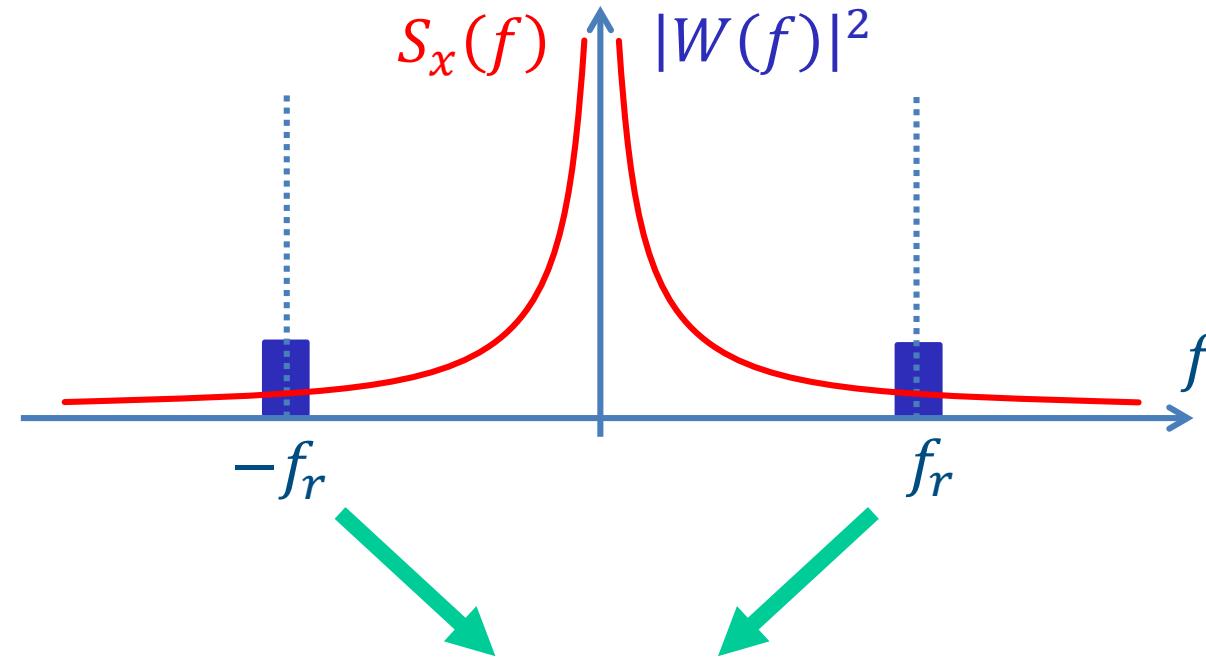
$$\overline{n_y^2} = \int S_d(f) |W_{LP}(f)|^2 df$$

$$= \frac{B^2}{4} \left( \int S_x(f - f_r) |W_{LP}(f)|^2 df + \int S_x(f + f_r) |W_{LP}(f)|^2 df \right)$$

$$= \frac{B^2}{4} \left( \int S_x(f) |W_{LP}(f + f_r)|^2 df + \int S_x(f) |W_{LP}(f - f_r)|^2 df \right)$$

$$= \int S_x(f) |W(f)|^2 df$$

# Output noise



These are the input noise components that will contribute to the output noise  $\Rightarrow$  remember the meaning of the weighting function!

# Output rms noise

- We consider the equivalent noise bandwidth of  $W_{LP}(f)$ ,  $BW_n$
- We approximate  $S_d(f)$  with  $S_d(0)$  over  $BW_n$  (narrow-band output filter)
- The output mean square noise is then

$$\overline{n_y^2} = \int S_d(f) |W_{LP}(f)|^2 df \approx S_d(0) 2BW_n$$

$$= 2BW_n \frac{B^2}{4} (S_x(-f_r) + S_x(f_r)) = B^2 BW_n S_x(f_r)$$

## Previous example:

- $B = 1 \text{ V}$ ,  $f_r = 1.5 \text{ kHz}$ ,  $T_F = 1 \text{ s} \Rightarrow BW_n = 0.25 \text{ Hz}$
- WN case:

$$\overline{n_y^2} = \lambda BW_n = 2.5 \times 10^{-12} \text{ V}^2 \Rightarrow \sqrt{\overline{n_y^2}} = 1.58 \mu\text{V}$$

- FN case:

$$\overline{n_y^2} = \frac{K}{f_r} BW_n = 5.5 \times 10^{-12} \text{ V}^2 \Rightarrow \sqrt{\overline{n_y^2}} = 2.35 \mu\text{V}$$

# S/N ratio

- We consider a constant (modulated) signal

$$x(t) = A \cos(\omega_r t)$$

- Output signal and noise are

$$y(t) = \frac{AB}{2} \quad (\text{no phase errors})$$

$$\overline{n_y^2} = B^2 BW_n S_x(f_r)$$

$$\frac{S}{N} = \frac{A}{\sqrt{4S_x(f_r)BW_n}}$$

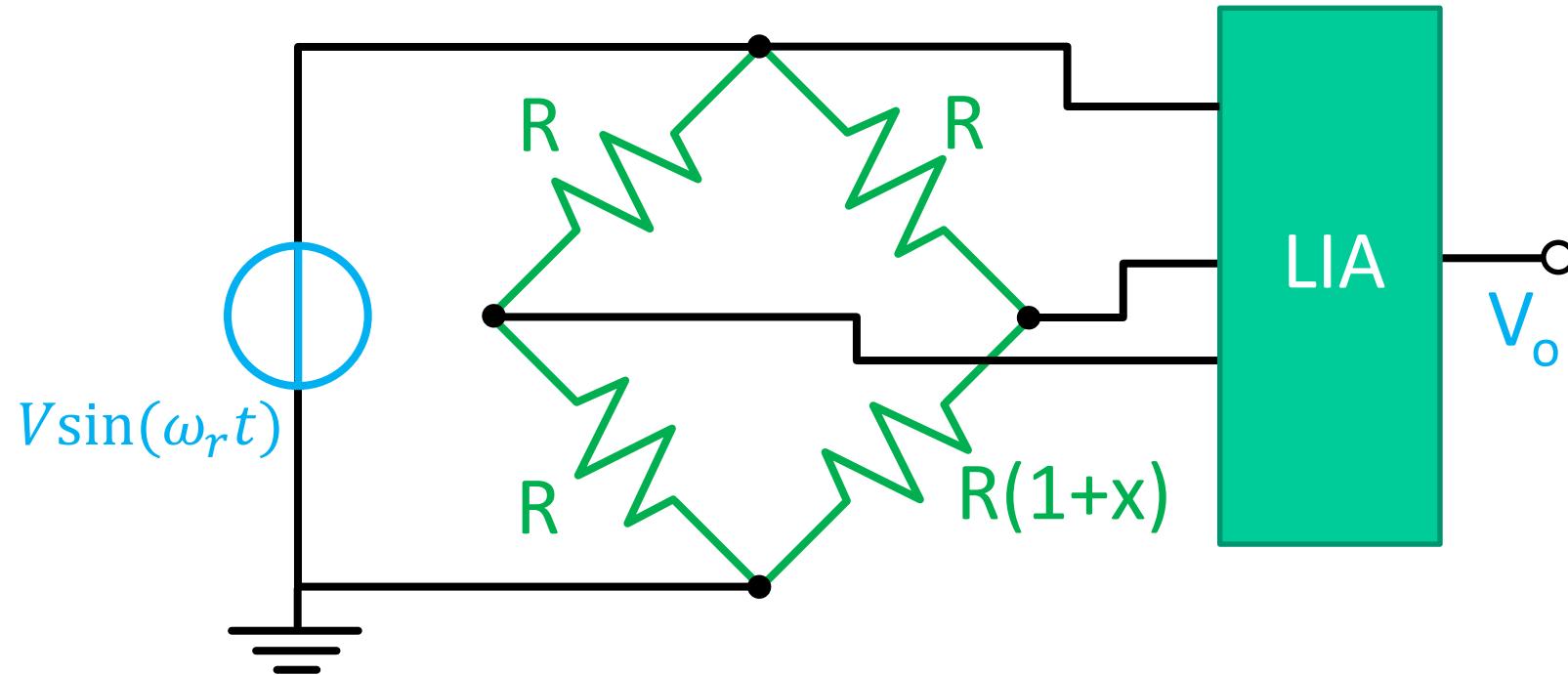
# Remarks

- If only LPF were used,  $S/N$  would be

$$\frac{S}{N} = \frac{A}{\sqrt{2S_x(0)BW_n}}$$

- Synchronous detection is useful only if  $S_x(0) \gg S_x(f_r)$
- The modulation stage must be inserted **before** the relevant LF noise sources (usually the amplifiers)

# Wheatstone bridge with LIA



# Low-light measurement with LIA

